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Wide-angle parabolic approximations for the nonlinear Helmholtz equation in the Kerr media

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Abstract – The method of multiple scales is used to derive parabolic approximations for the nonlinear Helmholtz equation in the Kerr medium. The resulting approximation has the form of a system of iterative parabolic equations. The zero-order approximation coincides with the solution of the standard nonlinear Schrödinger equation. High-order corrections are obtained by solving the linear parabolic equations with the input terms computed from the solutions of the previous equations of the iterative system. The developed approach is applied to the modeling of the nonparaxial optical soliton propagation. It is shown that the wide-angle parabolic approximations

allow to take nonparaxiality effects into account and eliminate the significant phase error produced

Introduction. – The propagation of continuous-wave 1 laser beam in the media featuring Kerr nonlinearity is de-2 scribed by the scalar nonlinear Helmholtz equation (NHE) 3 [1]. The development of various approaches to its solu-4 tion is very important in many practical contexts. Very few studies report far-reaching results related to the di-6 rect numerical solution techniques for the NHE [2,3]. The iterative methods proposed in the latter works are somewhat computationally intense and therefore they can pose g certain restrictions on the domain size in the applications. 10

by the nonlinear Schrödinger equation.

For many practical purposes the NHE can be approx-11 imated by the nonlinear Schrödinger equation (NSE) [1] 12 which is more computationally efficient. Being a one-way 13 paraxial approximation to the NHE, the NSE completely 14 neglects backscattering effects, and is inaccurate for the 15 modeling of nonparaxial propagation. The neglect of non-16 paraxial effects often manifests is the so-called phase er-17 ror that becomes considerable for the large propagation 18 distances. Elimination of constraints posed by the parax-19 ial approximation is a challenging task in various fields, 20 for instance, in optical communication [10,11], endoscopic 21 imaging using multimode lasers [12, 13], soliton-soliton 22 23 scaterring [14], multiplexing of non-dispersive beams [15] and solitonic turbulence [16–19]. Non-paraxial phenomena 24 also manifest in the modelling of intense light propagation 25

through randomly inhomogeneous active media [20].

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In the linear theory of wave propagation the nonparaxial effects are usually taken into account by the so-called wideangle parabolic approximations [4, 5]. The traditional approach [4] to the derivation of wide-angle parabolic equations (PEs) cannot be readily generalized to the case of the NHE, as it is based on the formal factorization of the Helmholtz operator and the subsequent approximation of the operator square root. Recently a new approach to the derivation of high-order parabolic approximations for the linear Helmholtz equation was proposed [5, 6]. It is based on the method of multiple scales [7] and results in an iterative system of PEs, where the solution of the *n*-th equation is used to compute the input term of the n+1-th equation.

In this work we show that the approach from [5] can 41 be easily generalized to the nonlinear case. The resulting 42 parabolic approximation also has the form of the system of 43 iterative PEs. The first equation in the system coincides 44 with the NSE (and it is natural to call it narrow-angle, or 45 paraxial nonlinear PE), while the high-order corrections 46 are still computed from the linear PEs (just as in linear 47 case). The study is concluded with an example where 48 the nonparaxial soliton solution [8,9] is compared against 49 its high-order parabolic approximations. It is shown that 50

⁵¹ our approach indeed allows to take the nonparaxial prop-

⁵² agation effects into account and eliminate the phase error ⁵³ produced by the NSE. Although our approach also ne-

⁵³ produced by the NSE. Although our approach also ne-⁵⁴ glects backscattering, it can be used as a compromise be-

glects backscattering, it can be used as a compromise be tween the accurate but costly direct methods of the NHE
 solution and very efficient NSE that does not take into

account many important propagation features (especially
 nonparaxiality).

Helmholtz equation in the Kerr medium. – Consider the Helmholtz equation in a nonlinear Kerr-type medium:

$$\frac{\partial^2}{\partial z^2}E + \frac{\partial^2}{\partial x^2}E + k_0^2(1+\epsilon|E|^{2\sigma})E = 0, \qquad (1)$$

where E = E(x, z) denotes the electric field. Hereafter we 59 restrict our attention to the case $\sigma = 1$ (though it will be 60 clear that our approach can be readily used for other pos-61 sible values of σ . In this section we derive an iterative sys-62 tem of PEs, whose solutions form a wide-angle parabolic 63 approximation for the NHE (1). Note that the derivation 64 here follows closely the case of the linear Helmholtz equa-65 tion [5], and the only difference arises from the nonlinear 66 term in (1). 67

Using the standard multiple-scale approach [7], we introduce the following slow variables (using the so-called 'parabolic scaling')

$$Z = \epsilon z, \qquad X = \epsilon^{1/2} x,$$

and the fast variable $\eta = (1/\epsilon)\theta(X, Z)$. Consequently, the derivatives in (1) are replaced according to the chain rule:

$$\frac{\partial}{\partial z} \to \epsilon \left(\frac{\partial}{\partial Z} + \frac{1}{\epsilon} \theta_Z \frac{\partial}{\partial \eta} \right) ,$$
$$\frac{\partial}{\partial x} \to \epsilon^{1/2} \left(\frac{\partial}{\partial X} + \frac{1}{\epsilon} \theta_X \frac{\partial}{\partial \eta} \right) ,$$

⁶⁸ where the subscripts denote partial derivatives.

Next we postulate the following asymptotic expansion

$$E(x,z) = \mathcal{E}(X,Z,\eta) = \mathcal{E}_0(X,Z,\eta) + \epsilon \mathcal{E}_1(X,Z,\eta) + \epsilon^2 \mathcal{E}_2(X,Z,\eta) + \dots$$
(2)

Introducing the asymptotic expansion (2) into the equation (1) and collecting the terms of order ϵ^{-1} , ϵ^{0} , etc, we obtain an infinite sequence of equalities.

There is only one term $\epsilon^{-1}\theta_X$ of the order of ϵ^{-1} , and we have $\theta_X = 0$, hence $\theta = \theta(Z)$, i.e. the fast scale only depends on the range.

Terms of the order ϵ^0 are combined into the following equation

$$(\theta_Z)^2 \mathcal{E}_{0\eta\eta} + k_0^2 \mathcal{E}_0 = 0 \,.$$

In order to satisfy this equality we simply put

$$(\theta_Z)^2 = k_0^2$$

and readily obtain

$$\mathcal{E}_0 = \exp(\mathrm{i}\eta)\mathcal{A}_0(X,Z) \,.$$

At this point we choose the branch $\theta_Z = k_0$ of the solution of the Hamilton-Jacobi equation (3) and thus retain only the waves propagating in the positive direction of the X-axis (this is exactly the point where the one-way approximation in applied is our approach). From (3) we also find that

$$\theta(Z) = k_0 Z$$

Note that the uniformity of the asymptotic expansion (2) is maintained if and only if

$$\mathcal{E}_j = \exp(\mathrm{i}\eta)\mathcal{A}_j(X,Z)\,,\tag{4}$$

for all $j \ge 1$ (this is a typical result of the application of the multiple-scale expansion method (see [5,7]).

Also note that in the light of representation of \mathcal{E}_j in (4) we immediately rewrite the nonlinear term in (1) as

$$k_0^2 \epsilon |\mathcal{E}|^2 \mathcal{E} = k_0^2 \epsilon \exp(i\eta) (\mathcal{A}_0 + \epsilon \mathcal{A}_1 + \dots)^* (\mathcal{A}_0 + \epsilon \mathcal{A}_1 + \dots)^2,$$
(5)

where f^* denotes the complex conjugate of f.

Since the exponentials can be cancelled in (1) for the terms of all orders in ϵ , we now rewrite the nonlinear part of the expression (5) in the following way

$$k_{0}^{2}\epsilon(|\mathcal{A}_{0}|^{2}\mathcal{A}_{0}) + k_{0}^{2}\epsilon^{2}(2|\mathcal{A}_{0}|^{2}\mathcal{A}_{1} + \mathcal{A}_{0}^{2}\mathcal{A}_{1}^{*}) + k_{0}^{2}\epsilon^{3}(2|\mathcal{A}_{0}|^{2}\mathcal{A}_{2} + \mathcal{A}_{0}^{2}\mathcal{A}_{2}^{*} + 2\mathcal{A}_{0}|\mathcal{A}_{1}|^{2} + \mathcal{A}_{0}^{*}\mathcal{A}_{1}^{2}) + \dots$$
(6)

Now we proceed to the terms of the positive orders in ϵ . Collecting the terms containing ϵ^1 we arrive at the following equality

$$2ik_0 \mathcal{A}_{0Z} + \mathcal{A}_{0XX} + k_0^2 \mathcal{A}_0 |\mathcal{A}_0|^2 = 0.$$
 (7)

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Equation (7) is basically the nonlinear (cubic) Schrödinger equation (NSE). The standard derivation of (7) that could be found in many works, e.g. [2,3]. In our view, the derivation presented here is somewhat more clear.

An important advantage of our multi-scale approach is the possibility to derive the higher order corrections to the NSE (7). These corrections may probably both improve the handling of the nonlinear effects and also describe more accurately wide-angle propagation. The latter capability was already validated in [5] for the linear Helmholtz equation. Collecting the terms of the order ϵ^{s+1} we obtain an equation for A_s :

$$2ik_0\mathcal{A}_{s,Z} + \mathcal{A}_{s,XX} + k_0^2 \left(\sum_{l+n+m=s} \mathcal{A}_l \mathcal{A}_n \mathcal{A}_m^*\right) + \mathcal{A}_{s-1,ZZ} = 0. \quad (8)$$

For each $s \geq 1$ equation (8) is a generalization of the linear Schrödinger equation with an input term $\mathcal{A}_{s-1,ZZ}$ that is computed from the solution of previous equation. Note (3) that the coefficients of (8) also contain $\mathcal{A}_{s-1}, \mathcal{A}_{s-2}, \ldots$

Approximate solution of the NHE (1). – Note that the equations (7), (8) are written in slow variables Z, X. Switching back to the physical variables, and recalling that $\eta = \epsilon^{-1} \theta(Z) = \epsilon^{-1} k_0 Z = k_0 z$, we find that the solution of (1) can be approximated by the truncated series of N terms

$$E(x,z) \sim \exp(ik_0 z) \sum_{j=0}^{N} A_j(x,z),$$
 (9)

where A_i satisfy the equations

. . . .

$$\begin{aligned} 2ik_0A_{0z} + A_{0xx} + \epsilon k_0^2 |A_0|^2 A_0 &= 0, \\ 2ik_0A_{1z} + A_{1xx} + \epsilon k_0^2 \left(2|A_0|^2 A_1 + A_0^2 A_1^* \right) + A_{0zz} &= 0, \\ 2ik_0A_{2z} + A_{2xx} + \epsilon k_0^2 \left(2|A_0|^2 A_2 + A_0^2 A_2^* \right) + \\ \epsilon k_0^2 \left(2|A_1|^2 A_0 + A_1^2 A_0^* \right) + A_{1zz} &= 0, \\ & \cdots \\ 2ik_0A_{s,z} + A_{s,xx} + \epsilon k_0^2 \left(2|A_0|^2 A_s + A_0^2 A_s^* \right) + \\ \epsilon k_0^2 \left(\sum_{\substack{l+n+m=s, \\ l,n,m$$

(10)

Equations (10) can be solved one by one to obtain the ap-86 proximation for the solution to (1), and this enables us to 87 call (10) an *iterative system* of equations. Hereafter the 88 right hand side of (9) is called *N*-th order (wide-angle) 89 parabolic approximation for the solution of the NHE (1). 90 We use the term *iterative parabolic approximation* to em-91 phasize the difference between our approach and classical 92 wide-angle linear PE theory. 93

An example: non-paraxial soliton propagation. -One of very few exact solutions of NHE was proposed in [8, 9]. It describes a non-paraxial soliton propagating in the Kerr media. The solution [8] is described by its envelope function $A(x,z) = E(x,z) \exp(-ik_0 z)$, which writes as

$$A(x,z) = \frac{\sigma}{k_0} \sqrt{\frac{2}{\epsilon}} \operatorname{sech}\left(\frac{\sigma k_0 x + \sigma v z}{\sqrt{k_0^2 + v^2}}\right) \times \exp\left(-ivx\sqrt{\frac{k_0^2 + \sigma^2}{k_0^2 + v^2}} + ik_0 z \left(-1 + \sqrt{\frac{k_0^2 + \sigma^2}{k_0^2 + v^2}}\right)\right).$$
(11)

In fact the equation (11) contains two independent parameters v and σ . In the notation from [8] they are expressed as

$$\sigma = \sqrt{2} \frac{\eta}{w_0} ,$$
$$v = \sqrt{2} \frac{V}{w_0} ,$$

amplitude parameter, and w_0 is related to the width of the The first few polynomials computed using (15) are pre-

nonparaxial soliton [8]. Note that the expression E(x, y) = $A(x, z) \exp(ikz)$ satisfies the NHE (1) exactly. We now want to reproduce the solution (11) using the iterative PE system (10). Note that for the small paraxial parameter

$$\kappa = \frac{1}{4\pi^2 n_0^2} \left(\frac{\lambda}{w_0}\right)^2$$

(i.e. when $\kappa \to 0$, and the wavelength is much smaller than the soliton width w_0) the solution (11) turns into the solution of the NSE that coincides with the equation for A_0 :

$$2ik_0A_{0z} + A_{0xx} + \epsilon k_0^2 |A_0|^2 A_0 = 0.$$

The latter solution writes as

$$A_0(x,z) = \frac{\sigma}{k_0} \sqrt{\frac{2}{\epsilon}} \operatorname{sech}\left(\sigma x + \frac{\sigma v}{k_0}z\right) \times \exp\left(-\mathrm{i}vx + \frac{\mathrm{i}}{2k_0}\left(\sigma^2 - v^2\right)z\right). \quad (12)$$

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It is easy to see the difference between the nonparaxial soliton solution (11) and its paraxial approximation (12). Even this relatively simple example clearly demonstrates that the accuracy of the NSE approximation (7) can be insufficient. By contrast, the use of iterative PEs (8) results in an approximation of substantially higher quality.

For simplicity we consider the propagation of the nonparaxial soliton with zero transverse velocity V = 0 using the exact solution (11) of the NHE as a reference. We solve the Cauchy problems for the equations (8) on the unbounded domain $\Omega = \{(x, z) \in \mathbb{R} \times \mathbb{R}^+\}$. The initial condition at z = 0 for the equation (7) is obtained from the formula (11):

$$A_0|_{z=0} = A(x,0) = \frac{\sigma}{k_0} \sqrt{\frac{2}{\epsilon}} \operatorname{sech}\left(\sigma x\right) , \qquad (13)$$

while zero Cauchy initial data $A_s|_{z=0} = 0$ at z = 0 for s > 00 is supplied to the high-order equations (8). Obviously, the solution of the Cauchy problem for (7) with the initial conditions (13) will therefore coincide with (12). Now our goal is to resolve the higher-order equations in (10). In this particular case this can be accomplished analytically. We introduce the following ansatz into the equations for A_1, A_2, \ldots :

$$A_s(x,z) = P_s(z)\operatorname{sech}(\sigma x) \exp\left(\frac{\mathrm{i}\sigma^2}{2k_0}z\right).$$
(14)

After some lengthy computation we find that the functions $P_{s}(z)$ are merely polynomials which can be obtained from a simple ordinary differential equation

$$\frac{dP_s}{dz} = \frac{\mathrm{i}}{2k} \frac{d^2 P_{s-1}}{dz^2} - \frac{\sigma^2}{2k^2} \frac{dP_{s-1}}{dz} - \frac{\mathrm{i}\sigma^4}{8k^3} P_{s-1} \,, \quad z > 0 \,, (15)$$

where V is the transverse velocity of the soliton, η is the subject to the obvious initial condition $P_s(0) = 0$ at z = 0.

sented here

$$P_{0}(z) = \sqrt{\frac{2}{\epsilon}} \frac{\sigma}{k},$$

$$P_{1}(z) = -\sqrt{\frac{2}{\epsilon}} \frac{i\sigma^{5}}{8k^{4}} z,$$

$$P_{2}(z) = -\sqrt{\frac{2}{\epsilon}} \frac{\sigma^{9}}{128k^{7}} z^{2} + \sqrt{\frac{2}{\epsilon}} \frac{i\sigma^{7}}{16k^{6}} z,$$

$$P_{3}(z) = \sqrt{\frac{2}{\epsilon}} \frac{i\sigma^{13}}{3072k^{10}} z^{3} + \sqrt{\frac{2}{\epsilon}} \frac{\sigma^{11}}{128k^{9}} z^{2} - \sqrt{\frac{2}{\epsilon}} \frac{5i\sigma^{9}}{128k^{8}} z^{3}.$$
(16)

Here we consider the evolution of the soliton with the 100 transverse velocity V = 0, energy parameter $\eta = 1$, 101 and width $w_0 = 1.005 \cdot 10^{-6}$ m, the light wavelength 102 in vacuum is $\lambda = 1 \,\mu$ m. The soliton is evolving in the 103 Kerr medium with refractive index $n_0 = 2.0$, and Kerr 104 nonlinearity parameter $n_2 = 0.2 \cdot 10^{-16} \text{ m}^2/\text{W}$ (in our 105 notation $\epsilon = 2n_2/n_0$). We compute the solution for 106 $z \in [0, 6 \cdot 10^{-3} \text{ m}]$. The comparison of the solution of 107 the NHE (1) and the respective parabolic approximations 108 (9) is presented in Fig. 1. Note that the improvement 109 due to the high-order terms A_s is noticeable even for 110 $z \in [1 \text{ mm}, 1.5 \text{ mm}]$ (see top subplot in Fig. 1), and the 111 phase error of the paraxial solution A_0 in clearly visible. 112 This phase error grows further for $z \in [3 \text{ mm}, 3.5 \text{ mm}]$ (see 113 middle subplot in Fig. 1), and we have to include more 114 terms of the iterative parabolic approximation in order to 115 obtain more accurate solution amplitude at greater ranges. 116 For $z \in [5 \text{ mm}, 5.5 \text{ mm}]$ (see bottom subplot in Fig. 1) we 117 find that the third-order iterative parabolic approxima-118 tion is still highly accurate, while lower-order approxima-119 tions fail to reproduce either correct phase or amplitude. 120 Overall, the advantage of the high-order parabolic approx-121 imations (9),(8) over the standard paraxial equation (7) is 122 clearly demonstrated in our example. 123

Conclusion. – In this study we presented the con-124 struction of the high-order parabolic approximations for 125 the NHE. On one hand, the derived system of the iter-126 ative PEs is a direct generalization of the parabolic ap-127 proximations obtained in the previous work for the linear 128 Helmholtz equation [5]. On the other hand, our theory 129 provides corrections to the NSE that allow to take the 130 nonparaxiality effects into account and eliminate the phase 131 errors produced by the NSE. There are still many ques-132 tions related to the parabolic approximations proposed 133 here, that should be resolved. In particular, it is neces-134 sary to develop efficient numerical methods, derive some 135 artificial boundary conditions (e.g. a generalization of the 136 boundary conditions from [6]) and obtain some energy es-137 timates similar to the asymptotic energy flux conservation 138 theorem [5]. We also stress that our approach does not 139 compete with the direct numerical methods for the NHE 140 solution, e.g. [2,3], as it is not capable of handling the 141 backscattering. Yet it presents a more computationally 142 efficient alternative which can be useful for applications, 143

where the nonparaxial propagation is dominant.

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Fig. 1: The real part of the exact solution A(x, z) of the NHE (solid line) and the parabolic approximations of the order 0 to 4 (dashed lines) for different intervals of z: $z \in [1 \text{ mm}, 1.5 \text{ mm}]$ (top subplot), $z \in [3 \text{ mm}, 3.5 \text{ mm}]$ (middle subplot), $z \in [5 \text{ mm}, 5.5 \text{ mm}]$ (bottom subplot).